

STATE TRANSITION MATRIX OF RELATIVE MOTION FOR THE NONCIRCULAR ORBIT. RELATION WITH PARTIAL-DERIVATIVE MATRIX IN THE SATELLITE COORDINATE SYSTEM

Andrey I. NAZARENKO
 Space Observation Center, Russia
nazarenko@iki.rssi.ru

ABSTRACT

The analytical formulas for the partial-derivative matrix $U(t, t_0)_{RSW}$ in the satellite coordinate system were derived by the Russian experts P.E. Elyasberg [1] and V.I. Charniy [2] at the beginning of sixties. This matrix includes the effect due to the reference orbit eccentricity. The relation between state vector variations at some time instant in geocentric celestial (x, y, z) and satellite coordinate system (R, S, W) is appeared by the orthogonal transformation matrix G

$$\begin{vmatrix} \frac{\partial(x, y, z, v_x, v_y, v_z)}{\partial(R, S, W, V_R, V_S, V_W)} \end{vmatrix} = \begin{vmatrix} G & 0 \\ 0 & G \end{vmatrix}$$

It is shown in this paper that the mentioned matrix $U(t, t_0)_{RSW}$ can be serve as the basis for calculating the state transition matrix of relative motion for a non-circular orbit. The derived fomulas

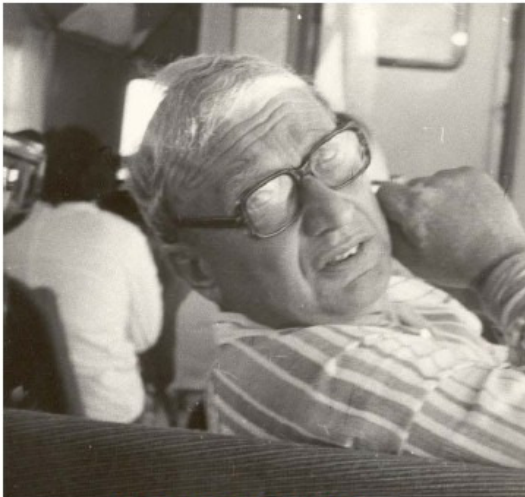
$$U(t, t_0)_{rel} = L(t) \cdot U(t, t_0)_{RSW} \cdot L(t_0)^{-1},$$

$$L(t) = \begin{vmatrix} E & 0 \\ -G^T \dot{G} & E \end{vmatrix},$$

being rather simple in appearance, represent generalization of the well known solution of the Clohessy-Wiltshire equations, suitable for circular orbits only.

This paper presents as well the results of the matrix $U(t, t_0)_{rel}$ calculation by using the developed method and their comparison with results of numerical integration and application of the solution of the Clohessy-Wiltshire equations for circular orbits.

1. RESULTS OF P.E. ELYASBERG



Pavel E. Elyasberg, 1914-1988

Apparently, in the Russian literature, the case under consideration was stated for the first time in the monograph by P.E. Elyasberg [1] and in the publication by V.I. Charniy [2]. The solution of the problem is presented in P.E. Elyasberg's monograph

(Section 10.11 «Partial-derivative matrix of the current characteristics under the initial motion data in the rectangular coordinate system (at $t=\text{const}$)») in the most complete form.

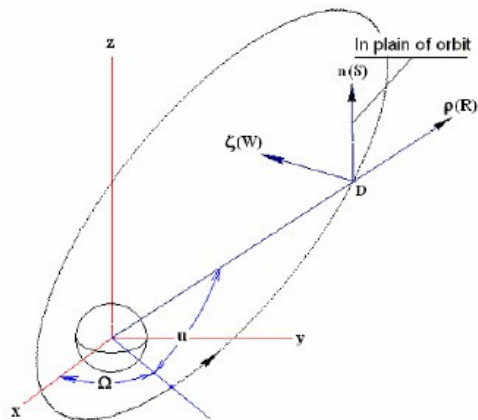


Figure 1. The orbital coordinate system

The technique and results of the partial-derivative matrix construction in the orbital coordinate system